

# AN INTRODUCTION TO THE NUMERICAL MODELING OF FUSION PLASMAS

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# Outline

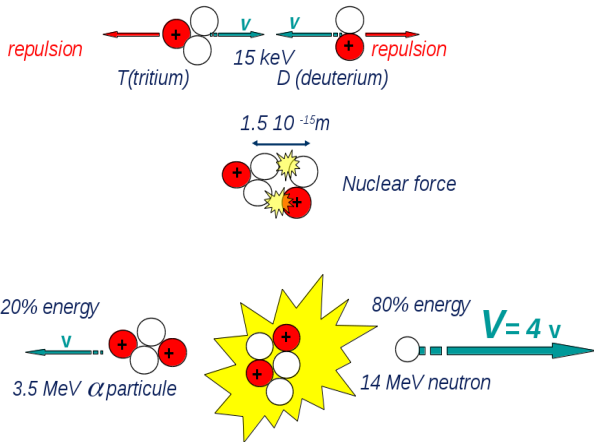
- 1 Fusion Plasmas
- 2 Kinetic Models
- 3 Fluid models
- 4 The MHD limit

# Outline

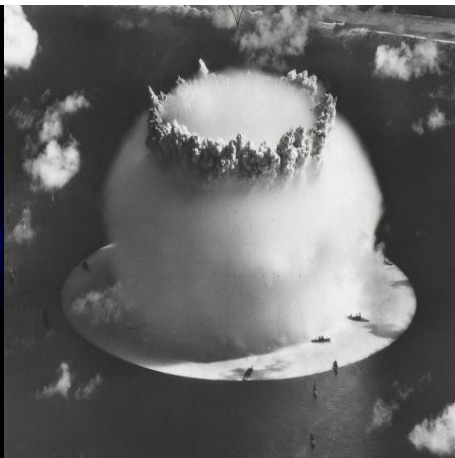
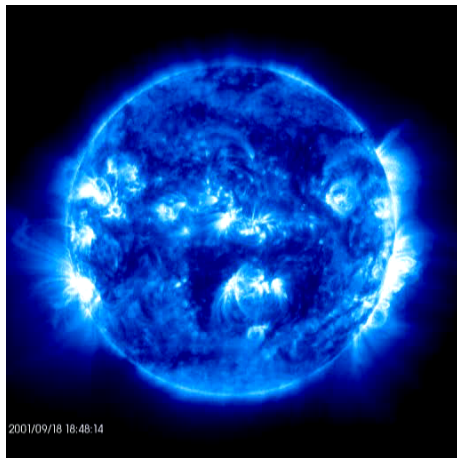
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# Deuterium-Tritium Reaction

## Deuterium tritium fusion



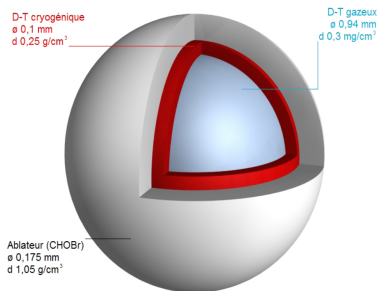
# Nuclear reactions



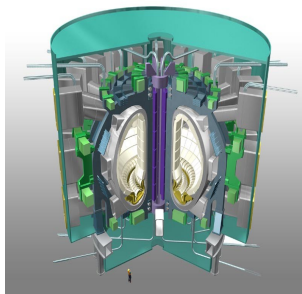
# Fusion on earth ???

Essentially two studied technologies :

## Inertial confinement

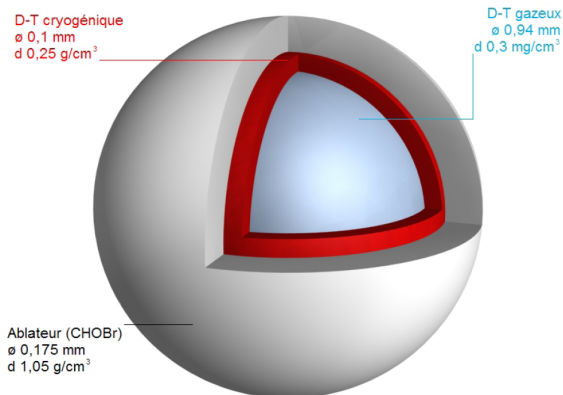


## Magnetic confinement



# Inertial confinement technology

relies on fast and violent heating of fusion targets.



# Target Ablator heating



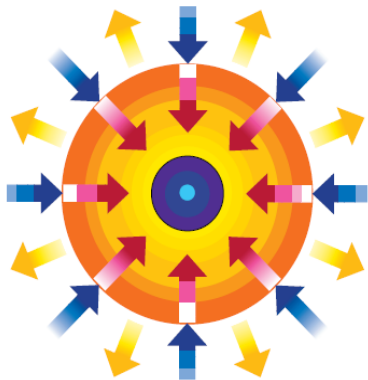
Ablator is illuminated by powerful - X rays and reaches a plasma state



# Ablator expansion & DT Compression

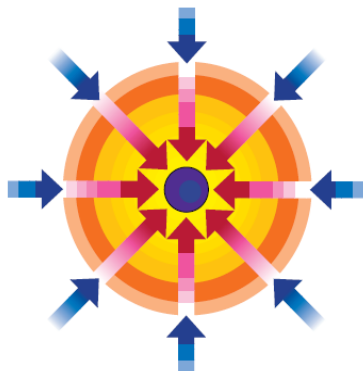
Ablator is accelerated outwards

By rocket effect DT is accelerated inwards ( $v \sim 300$  km/s)



# Deuterium - Tritium Compression

DT ice is compressed up to 1000 times its initial density ( 20 time lead density)



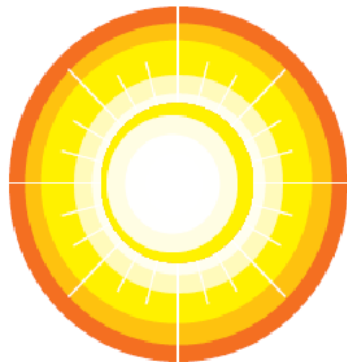
# Hot spot Ignition

Central gas is heated up to several  $10^6$  of degrees

When  $\rho R \sim 0.3\text{g}/\text{cm}^2$  the gas is self heated by alpha particles

Fusion reactions start and propagate into DT ice

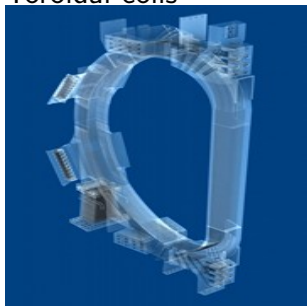
Note : 1 mg D-T  $\rightarrow$  340 MJ : Fusion is equivalent to combustion of 10 kg coal



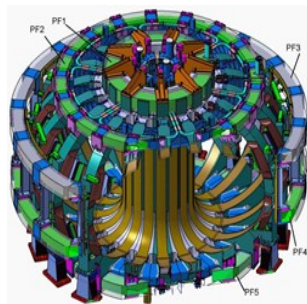
# The other way to fusion : magnetic confinement

Objective : confine hot D-T plasma by strong magnetic fields

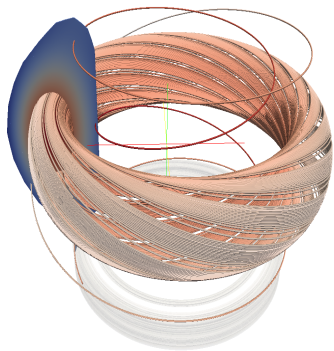
Toroidal coils



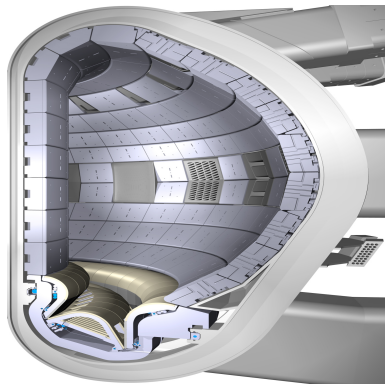
Poloidal coils



# Tokamaks :



magnetic field lines



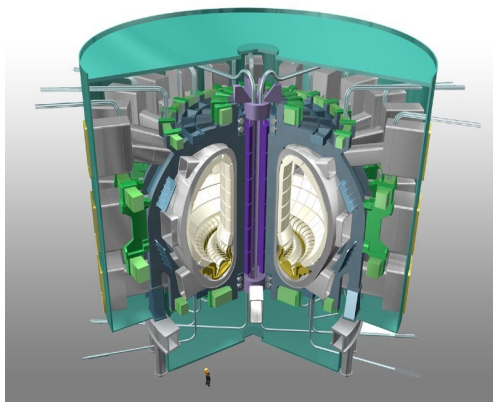
vacuum chamber

Tokamaks design in the 1950s by Igor Tamm and Andrei Sakharov

# ITER Tokamak

ITER (International Thermonuclear Experimental Reactor) in construction in Cadarache, construction begins in 2010, first plasma in 2020, first fusion plasma in 2027

- 840  $m^3$
- 150M  $^{\circ}K$
- 13 Tesla = 200 000 x earth magnetic field



# Numerical simulations

Necessary for

- Equilibrium computation : steady state and control of the machines
- prediction of the possible occurrence of instabilities
  - magnetic instabilities
  - hydrodynamic instabilities
- determination of the value of key parameters e.g transport coefficients due to turbulence
- understand and explain physical phenomena

Very large number of different numerical models

# Outline

- 1 Fusion Plasmas
- 2 Kinetic Models**
- 3 Fluid models
- 4 The MHD limit



# Kinetic models

Vlasov or Boltzmann eq for each particle (ions, electrons, neutral)

$$\frac{\partial f_s}{\partial t} + \operatorname{div}_x(\mathbf{v}f_s) + \operatorname{div}_v(\mathbf{F}f_s) = \mathbf{C}_s = \sum C_{ss'}$$

Force field  $\mathbf{F} = \frac{e_s}{m_s} [\mathbf{E} + \mathbf{v} \times \mathbf{B}]$  (note  $\operatorname{div}_v \mathbf{F} = 0$ )

$\mathbf{E}, \mathbf{B}$  given by Maxwell equations:

$$\frac{\partial \mathbf{B}}{\partial t} + \operatorname{curl} \mathbf{E} = 0 \quad \Upsilon = \sum e_s \int f_s(\mathbf{x}, \mathbf{v}, t) dv^3$$

$$-\frac{1}{c^2} \frac{\partial \mathbf{E}}{\partial t} + \operatorname{curl} \mathbf{B} = \mu_0 \mathbf{J} \quad \mathbf{J} = \sum e_s \int f_s \mathbf{v}(\mathbf{x}, \mathbf{v}, t) dv^3$$

$$\epsilon_0 \operatorname{div} \mathbf{E} = \Upsilon \quad \operatorname{div} \mathbf{B} = 0$$

# Kinetic models

- Self-consistent and closed model
- extremely heavy from computational point of view
  - 6 D model : 3 space dimensions, 3 velocity dimensions
  - covers huge range of time and space scales
- some simplifications possible : Ampere law, Quasi-Neutral assumption, Electrostatic assumption, gyrokinetic theory
- Used for some very specific tasks

# Kinetic models : one example of applications

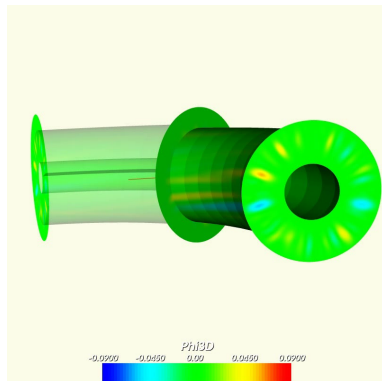
## Anomalous transport in tokamaks

Radial diffusion of mass and temperature exceed by order of magnitude “laminar” values

due to micro-turbulence and micro-instabilities :

Estimate the value of turbulent transport coefficients by direct simulation

Development of gyrokinetic codes (e.g GYSELA code of CEA)



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# From kinetic to fluid

Idea of fluid models :

instead of computing the whole distribution function  $f_s$   
compute only a small number of its moments, i.e : Fluid variables are moments of the distribution function  $f_s$

- fluid density

$$n_s(\mathbf{x}, t) = \int f_s(\mathbf{x}, \mathbf{v}, t) d\mathbf{v}^3$$

- fluid velocity

$$n_s(\mathbf{x}, t)\mathbf{u}_s(\mathbf{x}, t) = \int f_s(\mathbf{x}, \mathbf{v}, t)\mathbf{v}d\mathbf{v}^3$$

# From kinetic to fluid

- Pressure tensor

$$P_s(\mathbf{x}, t) = m_s \int f_s(\mathbf{x}, \mathbf{v}, t) (\mathbf{v} - \mathbf{u}_s) \otimes (\mathbf{v} - \mathbf{u}_s) dv^3$$

- scalar pressure = 1/3 trace of pressure tensor

$$p_s(\mathbf{x}, t) = \frac{m_s}{3} \int f_s(\mathbf{x}, \mathbf{v}, t) |\mathbf{v} - \mathbf{u}_s|^2 dv^3$$

- temperature

$$T_s(\mathbf{x}, t) = \frac{p_s(\mathbf{x}, t)}{n_s(\mathbf{x}, t)}$$

- Energy flux

$$\mathbf{Q}_s(\mathbf{x}, t) = \frac{m_s}{2} \int f_s(\mathbf{x}, \mathbf{v}, t) |\mathbf{v}|^2 \mathbf{v} dv^3$$

More complex fluid models, e.g 14 Moment closure by D. Levermore

# From kinetic to fluid

## Mass conservation

First velocity moment of the Boltzmann (Vlasov) equation

$$\int \left[ \frac{\partial f_s}{\partial t} + \operatorname{div}_x(\mathbf{v}f_s) + \operatorname{div}_v(\mathbf{F}f_s) \right] = C_s$$

gives

$$\int f_s dv^3 \stackrel{\text{def}}{=} n_s \quad \int \mathbf{v}f_s dv^3 \stackrel{\text{def}}{=} n_s \mathbf{u}_s$$

$$\int \operatorname{div}_v(\mathbf{F}f_s) = 0 \quad \int C_s dv^3 = 0$$

$$\frac{\partial n_s}{\partial t} + \operatorname{div}_x(n_s \mathbf{u}_s) = 0$$

proceeding in the same way for the other moments :

$m_s \mathbf{v} \times$  Boltzmann and integrating, gives momentum (velocity) equation

$\frac{1}{2} m_s |\mathbf{v}|^2 \times$  Boltzmann and integrating, gives energy (temperature) equation



$s \in \text{ion, electron, neutral particle}$

$$\frac{\partial n_s}{\partial t} + \operatorname{div}(n_s \mathbf{u}_s) = 0$$

$$m_s \left( \frac{\partial n_s \mathbf{u}_s}{\partial t} + \operatorname{div}(n \mathbf{u}_s \otimes \mathbf{u}_s) \right) + \nabla p_s + \operatorname{div}_x \Pi_s - n_s e_s (\mathbf{E} + \mathbf{u}_s \times \mathbf{B}) = \sum_{s' \neq s} R_{ss'}$$

$$\frac{\partial}{\partial t} \left( m_s \frac{n_s}{2} |\mathbf{u}_s|^2 + \frac{3}{2} p_s \right) + \operatorname{div} \mathbf{Q} - e_s \mathbf{u}_s \cdot \mathbf{E} = Q_s$$

# The 2-fluid model

- Much simpler than kinetic model (3D)
- Clear mathematical structure (1 compressible Navier-Stokes system per species coupled by force terms)
- But needs closure assumptions
- Questionable in tokamaks
- very large disparity in length and time scales
- still costly e.g for a ion-electron model 10 dof per mesh point
- not really used but good starting point for subsequent approximations

# From 2-fluid to one Fluid models

## Normalization units

time  $\tau_{obs}$

velocity  $u_*$

length  $L_* = u_* \tau_{obs}$

ion gyro-period  $\tau_{c_i} = m_i / Z_i e B$

electron gyro-period  $\tau_{c_e} = m_e / e B$

plasma period  $\tau_{pe} = (\epsilon_0 m_e / n_e e^2)^{1/2}$      $\tau_{c_e} / \tau_{pe} \sim 1$

collision  $\tau_{coll} \propto (\Lambda / \ln \Lambda) \tau_{pe}$

Non dimensional parameters

$$\varepsilon_i = \tau_{c_i} / \tau_{obs} \quad M_i = u_* / \nu_{T_i}, \quad \Lambda = \frac{4\pi}{3} n \lambda_D^3, \quad \mu = m_e / m_i$$

Relevant asymptotic regimes  $\varepsilon_i \rightarrow 0$      $\Lambda \rightarrow \infty$      $\mu = m_e / m_i \rightarrow 0$  etc

Give a huge set of one fluid models

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# From 2-fluid to one Fluid models

## The MHD scalings

MHD : violent instabilities affecting the whole plasma region :

$$x_* = e.g.a \text{ minor radius}$$

fast event : speed of MHD waves : Alfen velocity

$$v_A = \frac{B}{\sqrt{\mu_0 \rho}}$$

$$u_* = \mathcal{O}(v_{Ti}) \sim \mathcal{O}(v_A)$$

## MHD asymptotic limit

$$\varepsilon_i = \frac{\tau_{ci}}{\tau} \rightarrow 0, \quad M_i = \mathcal{O}(1)$$

# THE MHD MODELS

## Ideal MHD

Zero order model in  $\varepsilon_j$  ( $\varepsilon_j = 0$ )

$$\frac{\partial}{\partial t} \rho + \operatorname{div}(\rho \mathbf{u}) = 0$$

$$\frac{\partial}{\partial t} \rho \mathbf{u} + \operatorname{div}(\rho \mathbf{u} \otimes \mathbf{u}) + \nabla p - (\mathbf{J} \times \mathbf{B}) = 0$$

$$\mathbf{E} + \mathbf{u} \times \mathbf{B} = 0$$

need the current  $\mathbf{J} \rightarrow$  Ampere's law  $\mathbf{J} = \operatorname{curl} \mathbf{B}$

need the magnetic field  $\mathbf{B} \rightarrow$  Faraday's law

$$\frac{\partial \mathbf{B}}{\partial t} + \operatorname{curl} \mathbf{E} = 0$$

$\Rightarrow$  autonomous system for a **One Fluid model**

# THE MHD MODELS

## Ideal MHD - summary

$$\frac{\partial}{\partial t} \rho + \operatorname{div}(\rho \mathbf{u}) = 0$$

$$\frac{\partial}{\partial t} \rho \mathbf{u} + \operatorname{div}(\rho \mathbf{u} \otimes \mathbf{u}) + \nabla p + \frac{1}{\mu_0} (\mathbf{B} \times \operatorname{curl} \mathbf{B}) = 0$$

$$\frac{\partial \mathbf{B}}{\partial t} - \operatorname{curl}(\mathbf{u} \times \mathbf{B}) = 0$$

+ energy equation

# One example of hydrodynamic instabilities in magnetized plasmas

2 scalar variables : ion density  $n$ ; ion parallel velocity  $u_{\parallel}$

$$\mathbf{u} = u_{\parallel} \mathbf{b} + \mathbf{u}_{\perp}$$

$$\frac{\partial}{\partial t} \rho + \operatorname{div}(\rho \mathbf{u}) = 0$$

$$\frac{\partial}{\partial t} \rho u_{\parallel} + \operatorname{div} \rho u_{\parallel} \mathbf{u} + \nabla_{\parallel} p = \rho \mathbf{u} \cdot \frac{D\mathbf{b}}{Dt}$$

with (this is the ideal Ohms law solved for the  $\perp$  velocity)

$$\mathbf{u}_{\perp} = \frac{\mathbf{E} \times \mathbf{B}}{B^2}$$

$\mathbf{B}$  given,  $\mathbf{E}$  computed by the adiabatic assumption



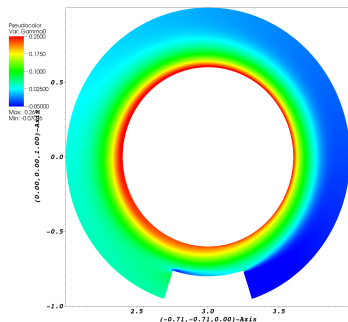
# Hydrodynamic instabilities

In the core plasma (white region)  
the plasma rotate counterclockwise  
In the edge region, the plasma  
touches the limiter

Due to electric charging of the wall  
(Bohm's BC) the plasma enters  
the wall

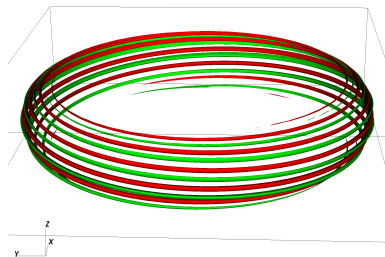
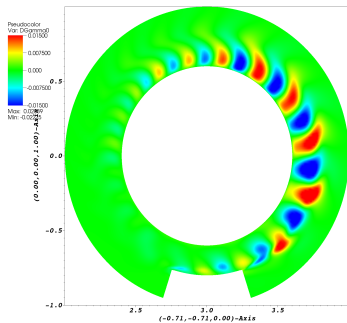
→ intense shear near the wall

Development of hydrodynamical in-  
stabilities (Kelvin-Helmholtz like) ?

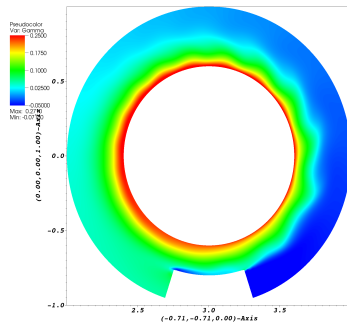


momentum plot at  $t=0$

# Hydrodynamic instabilities



Oscillation of the plasma boundary



# Conclusions

- large number of unsolved phenomena requiring the development of
  - numerical models
  - numerical methods
  - parallel algorithms
- short overview of numerical simulations for fusion plasmas
- See you next year in Cemracs2014